



# Exclusive Pair Production in ultra-relativistic Heavy Ion Collisions

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## 1 Equivalent Photon Method

- Nuclear Form Factor
- Photon Density
- Cross Section Calculation

## 2 Coherent $\rho^0$ -production

- Event- and Track Selection
- Invariant Mass Shape
- Efficiency Corrections
- Phase-Space MC Generator

## 3 Summary and Outlook

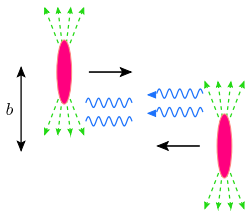
# Equivalent Photon Approximation



## Introduction

- E. Fermi (1924), Weizsäcker-Williams (1934)
- Idea: replace the electromagnetic field of a moving particle with an equivalent flux of photons.

$$\omega_L^{\max} = \frac{\hbar}{\Delta t} \sim \frac{\hbar c \gamma L}{b} \xrightarrow{b_{\min} = R_A} \frac{\hbar c \gamma L}{R_A}$$



$E_{CM}$ (GeV)	$\omega_L^{\max, \gamma\gamma}$ (GeV)
7000	200
3500	100
200	6

- C. Bertulani, S. Klein, J. Nystrand, Ann. Rev. Nucl. Part. Sci. 55, 271 (2005).
- S. Klein, E. Scannapieco, STAR Note 243 (1996).



# Equivalent Photon Method

## Nuclear Form Factor

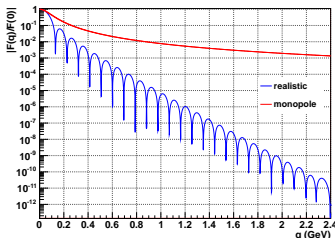
Nuclear charge density (Pb:  $R_A = 6.62$  fm,  $a = 0.546$  fm):

$$\rho(r) = \rho_0 \left[ 1 + \exp\left(\frac{r - R_A}{a}\right) \right]^{-1}$$

Nuclear form factors (Pb:  $\Lambda = 0.088$  GeV):

$$F(q) = \int_0^\infty \frac{4\pi}{q} \rho(r) \sin(qr) r dr$$

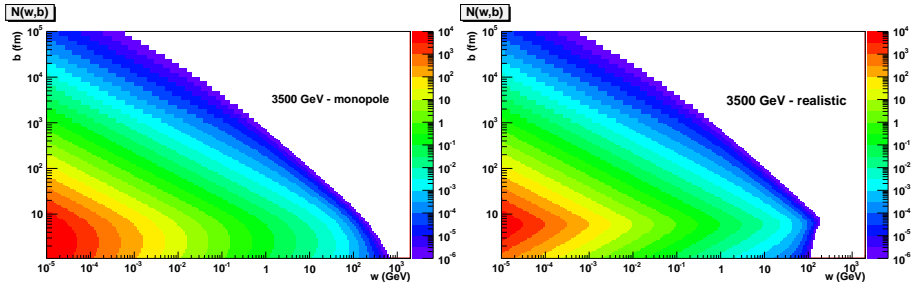
$$F^{\text{mon.}}(q^2) = \frac{\Lambda^2}{\Lambda^2 + q^2}$$



- S.R. Klein, J.Nystrand PRC 60(1999)014903, NPA 752(2005)470.
- M. Klusek-Gawenda and A. Szczurek et al. Phys. Rev. C 82 (2010) 014904, Int.J.Mod.Phys. A 26(2011) 741–743

# Photon density

Dependence on photon energy and on impact parameter



$$N_{\gamma}(w, b) = \frac{\alpha Z^2}{\pi^2 w b^2} \left[ \int_0^{\infty} du \frac{J_1(u)}{(bw/u\gamma)^2 + 1} F \left( \sqrt{\frac{u^2}{b^2} + \frac{w^2}{\gamma^2}} \right) \right]^2$$

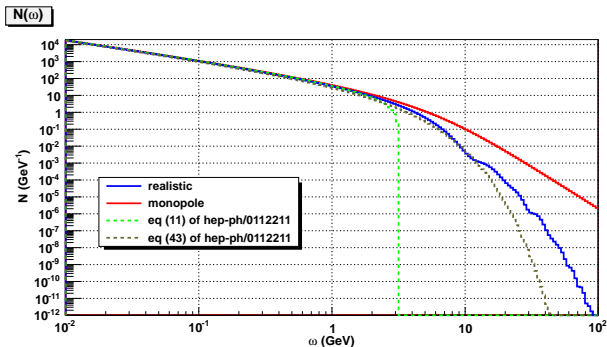
- For  $F^{\text{mon.}}$  the integral is numerically unstable  $\rightarrow$  use analytic result.
- For  $F^{\text{real.}}$ :  $u_{\text{max}} = b\sqrt{q_{\text{max}}^2 - (w/\gamma)^2}$ ,  $q_{\text{max}} = 2.4$



# Photon density

Dependence on impact parameter

$$N_\gamma(w) = \int d^2\vec{b} N_\gamma(w, b) \underset{w \text{ small}}{\sim} 2 \frac{\alpha Z^2}{\pi} \ln\left(\frac{\gamma}{wR_A}\right)$$



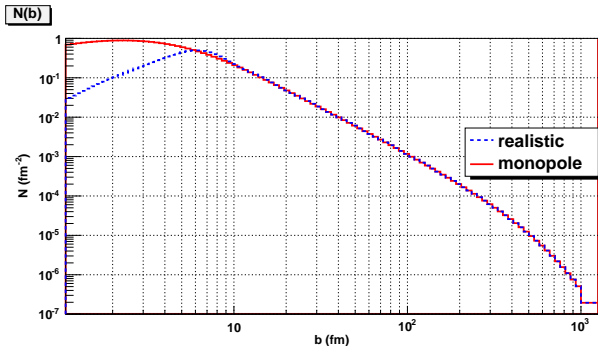
- For small  $w$  all four approaches agree



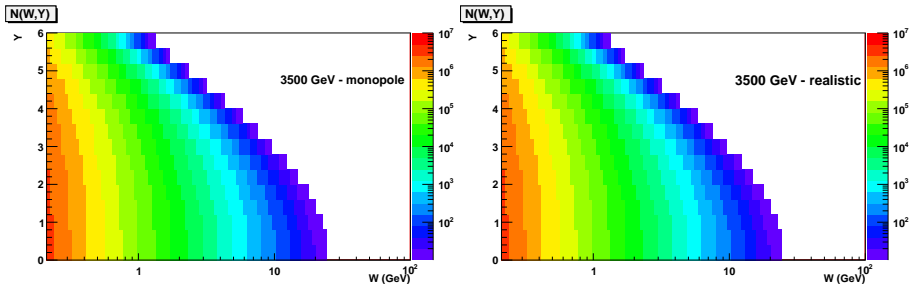
# Photon density

Dependence on photon energy

$$N_{\gamma}(b) = \int_0^{\infty} dw N_{\gamma}(w, b)$$



- at small impact parameters the realistic form factor drops off
- agreement at large  $b$



$W, Y$  –  $\gamma\gamma$  CM energy and rapidity:

$$N_{\gamma\gamma}(W, Y) = \int d^2\vec{b}_1 d^2\vec{b}_2 \theta(|\vec{b}_1 - \vec{b}_2| - 2R_A) N_\gamma\left(\frac{W}{2}e^Y, b_1\right) N_\gamma\left(\frac{W}{2}e^{-Y}, b_2\right)$$

- $\theta$ -function can be replaced by a smooth step function
- $N_\gamma$  can be computed using monopole or realistic form factors



# Cross Section Calculation



- $\gamma\gamma$  process  $AA \rightarrow AA X^+ X^-$ : (e.g.  $X = \pi, e, \mu$ )

$$\sigma(AA \rightarrow AA X^+ X^-) = \int dW \frac{d\sigma(\gamma\gamma \rightarrow X^+ X^-)}{dW} \int dY \frac{W}{2} N_{\gamma\gamma}(W, Y)$$

- $\gamma$ -pomeron process  $AA \rightarrow AA V$  ( $V = \rho^0, J/\psi, \Upsilon$ )

$$\sigma(AA \rightarrow AA V) = \int dw \frac{dN_\gamma}{dw} \sigma(\gamma A \rightarrow VA)$$

- Including nuclear breakup<sup>1</sup>

$$\sigma(AA \rightarrow A^* A^* V) = \int d^2b P^{\text{br}}(b) \int dw \frac{d^3 N_\gamma}{dw d^2b} \sigma(\gamma A \rightarrow VA) \theta(|\vec{b}| - R_A)$$

- ▶  $P^{\text{br}}$ : probability of nuclear breakup
- ▶  $\theta(|\vec{b}| - R_A)$  can be replaced by more realistic function

<sup>1</sup>A. Baltz, S.R. Klein and J. Nystrand, Phys.Rev.Lett. 89 (2002) 012301

# Coherent $\rho^0$ Production



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- 2 Coherent  $\rho^0$ -production
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  - Invariant Mass Shape
  - Efficiency Corrections
  - Phase-Space MC Generator

- 3 Summary and Outlook

# Coherent $\rho^0$ Production

## Introduction



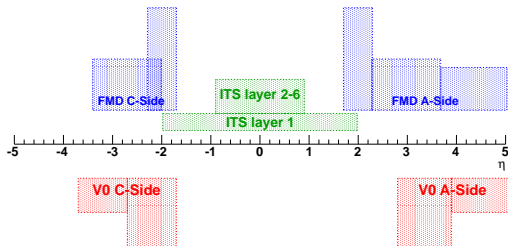
- $AA \rightarrow AA \pi^+ \pi^-$
- Branching ratio  $\rho^0 \rightarrow \pi^+ \pi^- \sim 100\%$
- Has been measured at RHIC<sup>2</sup>
- ALICE measurement, analysis is ongoing
- Event characterization:
  - ▶ two charged opposite-sign tracks at mid-rapidity
  - ▶ no tracks outside central rapidity
  - ▶ enhancement at low sum- $P_T$  (but track  $P_T$  can be big)
  - ▶ nuclear breakup: neutrons can be detected in zero-degree calorimeters

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<sup>2</sup>C. Adler et al. PRL 89(2002)272302

# Coherent $\rho^0$ Production

## Event- and Track Selection



- Ideal event: two charged tracks in an otherwise empty detector.
- Pseudo-rapidity gap, e.g. in ALICE veto on  $|\eta| > 0.9$ .
- Coherent production: low pair- $P_T$
- Event emptiness definition: detector noise, dependence on beam background
- Background estimation: like-sign events, analysis of minimum-bias MC.

# Coherent $\rho^0$ Production

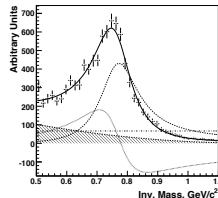
## Shape of Invariant mass distribution

The minv distribution is described by the Söding interference formula:

$$\frac{d\sigma}{dm_{\pi\pi}} = \left| A \frac{\sqrt{m_{\pi\pi} m_{\rho^0} \Gamma(m_{\pi\pi})}}{m_{\pi\pi}^2 - m_{\rho^0}^2 + im_{\rho^0} \Gamma(m_{\pi\pi})} + B \right|^2$$

with minv-dependent width:

$$\Gamma(m_{\pi\pi}) = \Gamma_{\rho^0} \frac{m_{\rho^0}}{m_{\pi\pi}} \left( \frac{m_{\pi\pi}^2 - 4m_{\pi}^2}{m_{\rho^0}^2 - 4m_{\pi}^2} \right)^{3/2}$$



- $B$  is the non-coherent *amplitude*
- $(A, B, \Gamma_{\rho^0}, m_{\rho^0})$  are determined by a fit.
- The ratio  $|B/A|$  ratio has been measured before, cf. e.g. by STAR<sup>3</sup>

<sup>3</sup>PRC 77 (2008) 034910

# Coherent $\rho^0$ Production

## Efficiency Corrections



- A real detector is not perfect: need to correct for efficiency  
     $\implies$  MC simulations.
- Goal: computing a corrected invariant mass histogram
- No unfolding is necessary ( $\Gamma_{\rho^0} \sim 150\text{MeV}$ )
- We use  $(1, N_Y, N_M)$  bins in  $P_T$ , rapidity and in invariant mass.
- The correction is applied for each bin in  $(Y, M)$ :

$$\left( \frac{d^2\sigma}{dM dY} \right)^{\text{abs}} = \frac{1}{\epsilon(M, Y)} \left( \frac{d^2\sigma}{dM dY} \right)^{\text{raw}}$$

- Used MC generators:
  - ▶ STARlight
  - ▶ Phase-space generator, see next slide

# Coherent $\rho^0$ Production

## Phase-space MC Generator



- Physics input:
  - ▶ kinematics  $1 \rightarrow 2$
  - ▶  $\theta$ -distribution of decaying particle  $\sim \sin^3 \theta$  ( $\rho^0 \rightarrow \pi^+ \pi^-$ )
- All remaining free variables are drawn from flat distributions
- Method is based on STARlight

- ① Generate the 4-vector of  $\rho^0$  with given  $M, Y, P_T$ : ( $\alpha \in [0, 2\pi]$ )

$$v = (P_T \cos \alpha, P_T \sin \alpha, M_T \sinh Y, M_T \cosh Y), \quad M_T = \sqrt{M^2 + P_T^2}$$

- ② Generate two pion-tracks in the rest frame of  $\rho^0$

$$v_{\pm} = (\pm \vec{p}, M/2) \quad \text{with} \quad p^2 = M^2/4 - m_{\pi}^2 \quad \text{and} \quad \theta(\vec{p}) \sim \sin^3 \theta$$

- ③ Then boost  $v_{\pm}$  from the  $\gamma\gamma$  CM frame into the laboratory frame using  $v$ .



- Equivalent Photon Approximation:
  - ▶ tool to compute UPC cross sections
  - ▶ different options for the nuclear form factor, shadowing
  - ▶ can be applied to nuclear breakup
- Ultra-peripheral collisions are characterized by
  - ▶ two charged opposite-sign tracks at mid-rapidity
  - ▶ low sum- $P_T$  (coherent production)
- ALICE is working on UPC processes  $PbPb$ 
  - ▶  $\rho^0 \rightarrow \pi^+\pi^-$
  - ▶  $J/\psi \rightarrow e^+e^-, \mu^+\mu^-$
- Absolute cross sections at mid-rapidity (work is ongoing)
- Nuclear breakup can be studied using zero-degree calorimeters
- Outlook: pPb collisions in 2012.





Thank you for your attention